Computational study of reaction between prebiotic molecules on surfaces

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Aim of this work is the computational modeling of the reaction between carbon monoxide and the prebiotic species C_nN (n = 1, 2, 3) in a carbon-rich dense molecular cloud which produces molecules that have all the elements of the amino acids backbone. The mentioned reactions have been investigated on two carbon surfaces: graphene, as a model of an interstellar grain covered by carbonaceous materials, and a polycyclic aromatic hydrocarbon, abundantly present in dense molecular clouds. [1-3] Among all the reactions investigated, the one leading to the C_3NO formation is by far the most exoergic, therefore it was extensively studied and two reaction mechanisms have been suggested. The first one is a Langmuir-Hinshelwood-type, investigated by means of DFT calculations. The second one is an Eley-Rideal-type, studied by *ab initio* molecular dynamics, performed at the DFT level according to the SIESTA method. The energetic aspects of the Eley-Rideal mechanism are detailed through a velocity distribution function (VDF) analysis.

Geometry of prebiotic molecules on surfaces

The geometries of C_nN and C₂O are reported on figure (a,b,c,d). The latter was studied in order to suggest a different reaction pattern for the formation of C_3NO_2 .



Prebiotic species adsorption produce a small distortion on the surface. In presence of a structural defect, the C1-C2 bonding is elongated by 0.10 Å, making the $-C_2N$ fragment activate. The adsorption of C_2O is the most exoergic, due to the bridge interaction. Carbon monoxide is phisisorbed on both carbonaceous surfaces.

Computational Details

The pseudopotentials used for the SIESTA method were generated for all atoms involved in the investigation. The Troullier-Martins parametrization was adopted and the semilocal form produced was transformed into the fully non local form proposed by Kleiman and Bylander. The pseudopotential generation process mainly involved three steps:



SIESTA*

DRSLL

-140.62

-80.52

-159.67

-31.72

-113.50

CN/g

C₂N / g

C₃N / g

CO/g

C₂O / g

M062X/cc-pvDZ*

ΔSCF

-169.50

-71.52

-12.28







1.16 1.20

1.38 1.39

1.21 1.24

The adsorption of CO in presence of a $C_n N$ molecule is an exoergic process, for all the prebiotic species. Co-adsorption geometry of CN in presence of C_2O converges at the product C₃NO one. The geometries of coadsorption of C_nN and CO are almost the same of those corresponding to the adsorption of the single species.

(KJ/mol)	SIESTA* DRSLL	M062X/c	c-pvDZ*
		ΔSCF	ΔΖΡΕ
CO + CN/g	-31.56	-6.75	-2.54
CO + C ₂ N/g	-27.21	-16.30	-11.53
$CN + C_2O/g$	-410.52	-407.17	-398.44
CO + C₃N/g	-34.23	-4.17	0.07

Reaction product – LH and ER mechanisms

1) generation of DFT atomic levels for the selected atom.

- 2) generation of the pseudopotential
- 3) check of the results.

We tested the pseudopotential in two steps: we calculated the difference between the atomic eingevalues obtained with the all-electron calculations and those computed with the pseudopotential (a value of 0.01 Ry for this difference was retained optimal). Then we performed a test calculation for the transferability, to evaluate the behavior of the pseudopotential in a chemical environment different from an atom.

Strictly localized basis sets where optimized using the pseudopotential previously generated. Our optimized basis set are equivalent to a double zeta polarized, considered accurate in this theoretical context. We exploited SIMPLEX algorithm in order to optimized variationally all the parameters defining the basis set. The standard of acceptance of basis set were the same applied for the transferability test of pseudopotential.

In the *ab initio* molecular dynamic we used the PBE functional.

In order to evaluate correctly the van der Waals interactions we used the M06-2X functional for the investigation on PAH with the Gaussian 09 software and the DRSLL functional with the SIESTA method on graphene.



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LH	SIESTA* DRSLL	M062X/o	cc-pvDZ*
		ΔSCF	ΔΖΡΕ
C(CO)N/g	-0.17	4.76	10.17
C(CO)CN/g	-212.11	-253.71	-238.36
CC(CO)N/g	-104.00	-163.76	-149.89
C(CO)CCN/g	2.183	-5.51	4.55
CC(CO)CN/g	5.64	-4.52	5.09
CCC(CO)N/g	129.16	119.30	127.23
ER	SIESTA* DRSLL	M062X/cc-pvDZ*	
		ΔSCF	ΔΖΡΕ
C(CO)N/g	-31.73	-1.99	7.63
C(CO)CN/g	-239.32	-270.01	-249.88
CC(CO)N/g	-131.21	-180.06	-161.42
C(CO)CCN/g	-32.05	-9.68	4.62
CC(CO)CN/g	-28.59	-8.69	5.16
CCC(CO)N/g	94.93	115.12	127.29



Astronomical complex organic molecules in different environments, 1st Italian workshop on astrochemisty - Florence 10-11 March 2016